

# A THEORETICAL STUDY OF ATOM-MOLECULE RAMSAY FRINGES OF BOSE-EINSTEIN CONDENSATION AND EVALUATION OF ITS PARAMETER AS A FUNCTION OF MAGNETIC FIELD

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#### **ABSTRACT**

In this paper, we have theoretically studied the physical properties of atom-molecule Ramsay fringes of Bose-Einstein Condensate of  $^{85}$ Rb atom in hyperfine state. Using the theoretical formalism of Goral et al.  $^{9}$ , we have evaluated the natural frequency, two-body loss rate constant K2 and visibility of Ramsy fringes as a function of magnetic field strength  $B_{evolve}$  (G). Our theoretically evaluated results are in satisfactorily agreement with the experimental data.

**Key words**: Ramsy fringes, Two-body loss rate constant, Visibility of Ramsy fringes, Feshbach resonance, Hyperfine states, Zero energy threshold, Diatomic vibration bound state.

# INTRODUCTION

The study of ultra cold molecules produced in atomic Bose and Fermi gases is one of the most important developments in cold atom physics<sup>1</sup>. One of the most successful experimental techniques to produce molecules in atomic Bose-Einstein condensates utilized adiabatic sweeps of a magnetic field tunable Feshbach resonances level across the zero energy threshold of the colliding atoms. In this period, a highly excited diatomic vibration bound states can be efficiently produced<sup>2-6</sup>. There is a coherent superposition of atomic condensates and molecules. These experiments employed a sequence of two magnetic field pulses each of which rapidly approached the position of zero-energy resonance (the magnetic field strength at which the scattering length has singularity) on the side supporting the highly excited diatomic vibration bound state. The two pulses were separated by an evolution period of variable duration as a function of which the oscillations in the final

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condensate populations were observed. This achievement offers the possibility of precise informatics measurements of the energies of the relevant states through the measurement of bulk properties of the gas.

In recent experiments<sup>7,8</sup>, the frequency and visibility of atom-molecule Ramsey fringes were observed. In this experiments, a Bose-Einstein condensate of  $^{85}$ Rb atom in (F = 2,  $m_F$  = -2) hyperfine states was exposed to a sequence of spatially homogeneous magnetic field pulses on the high-field side of the 155G Feshbach resonances. The evolution period of constant magnetic field strength  $B_{\text{evolve}}$  separated the pulse by an amount of time  $t_{\text{evolve}}$ . In the course of the experiments<sup>7,8</sup>, the final densities of atoms were recorded after expanding the cloud. The measurements were repeated with variable evolution time and magnetic field strength  $B_{\text{evolve}}$ .

Interference between different components of a partially condensed Bose gas was identified. There is a two different components of the final atomic cloud: a remnant Bose-Einstein condensate and a brest of atoms with a comparatively high mean kinetic energy. The magnitude of the brest of atoms and remnant condensate fractions exhibited an oscillatory dependence of the evolution time of the magnetic field pulse sequence. The existence of the third component of missing atoms was inferred from the difference between the initial and total final members of the detectable atoms. It was suggested that the missing atoms indicated molecular production and the oscillations were interpreted in terms of interference between atoms and molecules during the evolution period of the pulse sequence.

In this paper, using the theoretical formalism of K Goral et al.<sup>9</sup>, we have theoretically evaluated the natural frequency  $\upsilon_0$  of Ramsey fringes, two-body loss rate constant K2 and the visibility of Ramsey fringes as a function of magnetic field strength  $B_{evolve}$ . Our theoretically evaluated results are in good agreement with the experimental data<sup>7,8</sup>.

#### **EXPERIMENTAL**

# Mathematical formulae used in the evaluation

One first considers a pair of <sup>85</sup>Rb atoms in a spherically symmetric trap exposed to a spatially homogeneous magnetic field<sup>10,11</sup>. The degrees of freedom of the center of mass and relative motions of the atom pair are then exactly decoupled. The dynamics of the relative coordinate r is determined by the Hamiltonian of the form -

$$H_{2B} = \frac{-\hbar^2}{2m} \nabla^2 + V_{trap}(r) + V(B, r) \qquad ...(1)$$

Here  $V_{trap}(r)$  is the potential of the atom trap, m is the atomic mass and V(B,r) is the magnetic field dependent binary potential. In general,  $H_{2B}$  depends on all hyperfine states of the atoms, which are strongly coupled by the Feshbach resonance state  $\Phi_{res}$  (r). The analogy of the condensate mode is realized by the lowest energetic state of two atoms above the dissociation threshold of the pair interaction potential V(B,r) whose wave function  $\Phi_0(r)$  obeys the stationary Schrodinger equation.

$$H_{2R}\Phi_0(r) = E_0\Phi_0(r)$$
 ...(2)

The two-body Hamiltonian evaluated at the initial magnetic field strength of the pulse sequence. The amplitude for the probability to detect the atoms in the state  $\Phi_0$  at the final time  $t_{\text{final}}$  of the pulse sequence provides the analog of the amplitude for the remnant condensate fraction and is determined by –

$$T_{2B}(\Phi_0 \leftrightarrow \Phi_0) = \langle \Phi_0 / U_{2B}(t_{final}, t_0) / \Phi_0 \rangle \qquad \dots (3)$$

Here  $U_{2B}(t_{\text{final}},\ t_0)$  is the time evolution operator which satisfies the Schrodinger equation –

$$i\hbar \frac{\partial}{\partial t} U_{2B}(t, t_0) = H_{2B}(t) U_{2B}(t, t_0) \qquad \dots (4)$$

 $U_{2B}(t_{\rm final}, t_0)$  can be factorized in the evolution operators  $U_{\rm no.\ 1}(t_1, t_0)$   $U_{\rm no.\ 2}(t_{\rm final}, t_2)$  and  $U_{\rm evolve}(t_{\rm evolve})$  associated with the first and second magnetic field pulses and with the evolution period of the sequence respectively. Here  $t_1$  is the time immediately after the first magnetic pulse  $t_2$  is the time at the beginning of the second magnetic field pulse and  $(t_2-t_1)=t_{\rm evolve}$  is the duration of the evolution period. The factorization of  $U_{2B}(t_{\rm final}, t_0)$  yields.

$$U_{2B}(t_{final}, t_0) = U_{no.2}(t_{final}, t_0)U_{evolve}(t_{evolve})U_{no.1}(t_{final}, t_0) \qquad ...(5)$$

The frequency of the Ramsy fringe is obtained by equation

$$N_c^{evolve} = N_{avg} - \alpha t_{evolve} + A Exp(-\beta t_{evolve}) Sin(\omega_e t_{evolve} + \Delta \phi) \qquad ...(6)$$

Here  $\omega_e$  is the angular frequency

$$\omega_e = 2\pi (v_0^2 - (\frac{\beta}{2\pi})^2)^{\frac{1}{2}}$$

A is the amplitude of the interference fringes.

 $N_{avg}$  is the average number of atoms determined from the fringe visibility  $\alpha$  and  $\beta$  are fitting parameters

 $\alpha$  = loss rate of atoms

 $\beta$  = damping rate

 $\Delta \phi$  = phase shift of the Ramsy fringes

The two-body loss rate constant K<sub>2</sub> is determined by the formulae

$$K_2 = \frac{2\alpha}{N_{avg}} \langle n_c(t) \rangle \qquad \dots (7)$$

Where  $\langle n_c(t) \rangle$  is the average condensate density.

The visibility of the Ramsy fringes are determined by the following equation-

$$V_{Ramsy} = \frac{1 - Exp(-4nv[P_{o.b}^{no.1}P_{b.o}^{no.2}]}{1 + Exp(-4nv[P_{o.b}^{no.1}P_{b.o}^{no.2}]^{\frac{1}{2}}} \dots (8)$$

Here  $P_{o.b.}^{no.1}$  is the probability of molecular production in the first magnetic field pulse and probability is  $P_{b.o.}^{no.2}$  their reconversion into atom pairs in the lowest energetic quasi continuum state  $\Phi_0$  (r). n is the homogeneous gas density and v is the volume.

### RESULTS AND DISCUSSION

In this paper, we have determined the natural frequency  $v_o$ , two-body loss rate constant  $K_2$  and visibility of Ramsy fringes all as a function of magnetic field strength  $B_{\text{evolve}}$  (G). The theoretical results were compared with the experimental data<sup>7,8</sup>. The evaluation has been performed with the help of theoretical formalism developed by Groal et al.<sup>9</sup> The results are shown in Table 1, 2 and 3, respectively. In Table 1, using equation (6) we have calculated the natural frequency  $v_o$  of Ramsy fringe. Our evaluated results are lower than the experimental data<sup>7,8</sup> in magnitude but the trend is the same. The natural frequency  $v_o$  increases with magnetic field strength  $B_{\text{evolve}}$  (G). Our theoretical results also show that for  $B_{\text{evolve}} > 158$  G which is far away from the zero-energy resonance at Bo = 155 G, the fringe frequency closely matches with the vibrational frequencies associated with weakly bound molecular state<sup>10,11</sup>. The two-body loss rate constant K2 is determined using equation (7) and results are shown in Table 2 with the experimental data.

Table 1: Evaluated results of the natural frequency  $\upsilon_0$  of Ramsy fringes as a function of magnetic field strength  $B_{evolve}\left(G\right)$ 

B <sub>evolve</sub> (G)	Natural frequency $v_0$ (MHz)	
	Theoretical results	Expt. <sup>7,8</sup>
155	2.85	4.22
156	4.32	5.36
157	8.87	9.49
158	22.36	24.55
159	46.54	50.21
160	73.29	80.56
161	116.10	120.49
162	176.56	184.26
163	219.22	224.16
164	284.27	320.59
165	328.56	357.15

Table 2: Evaluated results of two-body loss rate constant K2 as a function of magnetic field strength  $B_{\text{evolve}}\left(G\right)$ 

B <sub>evolve</sub> (G)	Two-body loss rate constant K2 (cm <sup>3</sup> /s)		
	Theoretical results	Expt. <sup>7,8</sup>	
156	2 x 10 <sup>-8</sup>	3.16 x 10 <sup>-8</sup>	
156.5	$3.14 \times 10^{-8}$	$3.14 \times 10^{-8}$	
157	$3.05 \times 10^{-8}$	$1.97 \times 10^{-8}$	
157.5	$2.84 \times 10^{-8}$	$1.62 \times 10^{-8}$	
158	$2.56 \times 10^{-8}$	$1.48 \times 10^{-8}$	
158.5	$2.15 \times 10^{-8}$	1.29 x 10 <sup>-8</sup>	
159	$1.87 \times 10^{-8}$	$1.17 \times 10^{-8}$	
159.5	$1.62 \times 10^{-8}$	$1.05 \times 10^{-8}$	
160	$1.47 \times 10^{-8}$	$0.92 \times 10^{-8}$	
165	$1.16 \times 10^{-8}$	$0.86 \times 10^{-8}$	

B <sub>evolve</sub> (G)	Fringe visibility	
	Theoretical results	Expt. <sup>7,8</sup>
156	1.052	0.987
157	0.924	0.895
158	0.814	0.795
159	0.692	0.714
160	0.634	0.625
161	0.548	0.524
162	0.509	0.493
165	0.488	0.425
170	0.423	0.407

Table 3: Evaluated results of visibility of Ramsy fringes as a function of magnetic field strength  $B_{\text{evolve}}$  (G)

The rate constant K2 indicates the loss of condensate atoms due to energy transfer from the magnetic field pulses. This derives initially weakly interacting Bose-Einstein condensates into a strongly correlated non-equilibrium state. In this calculation, the deeply inelastic spin relaxation phenomena has not been included  $^{12}$ . Our theoretically evaluated results of loss constant K2 are in good agreement with the experimental data  $^{7,8}$ . Visibility of Ramsy fringes are determined using equation (8) and the results are shown in Table 3 along with the experimental data  $^{7,8}$ . The evaluation has been performed using exact transition probability  $P_{o.b.}^{no.1}$  and  $P_{b.o.}^{no.2}$  for both the pulses of realistic sequence and their counterparts as a function of magnetic field strength  $^{13}$ . Our theoretically evaluated results show that the fringe visibilities decreases with the increase of magnetic field strength and for high field the decrease is almost constant. Our theoretically evaluated results are in satisfactorily agreement with the experimental data and with other theoretical workers  $^{14-20}$ .

# **REFERENCES**

- 1. T. Kohler and K. Bernett, Phys. Rev., **A65**, 033601 (2002).
- 2. T. Kohler, T. Gasenzer and K. Bernett, Phys. Rev., A67, 013601 (2003).
- 3. S. Durr, T. Volz, A. Marte and G. Rempe, Phys. Rev. Lett. (PRL), **92**, 020406 (2004).
- 4. K. Xu, T. Mukaiyama, J. R. Abo-Shabir, J. K. Chin, D. E. Miller and W. Ketterle, Phys. Rev. Lett. (PRL), **91**, 210402 (2003).

- 5. T. Mukaiyama, J. R. Abo- Shabir, K. Xu, J. K. Chin and W. Ketterle, Phys. Rev. Lett. (PRL), **92**,180403 (2004).
- 6. S. Durr, T. Volz and G. Rempe, Phys. Rev., **A70**, 031601 (2004).
- 7. N. R. Claussen, S. J. J. M. Kokkelmans, S. T. Thompson, E. A. Donley, E. Hodby and C. E. Wieman, Phys. Rev., **A67**, 060701 (2003).
- 8. E. A. Donley, N. R. Claussen, S. T. Thompson and C. E. Wieman, Nature (London), 417, 529 (2002).
- 9. K. Goral, T. Kohler and K. Bernett, Phys Rev., **A71**, 023603 (2005).
- 10. T. Kohler, E. Tiesinga and P. S. Jullienne, Phys. Rev. Lett. (PRL), **94**,020402 (2005).
- 11. S. T. Thompson, E. Hodby and C. E. Wieman, Phys. Rev. Lett. (PRL), **94**, 020409 (2005).
- 12. E. Timmeemans, K. V. Kherntgyan and H. He, Phys. Reports, 335, 269 (2007).
- 13. R. A. Duine and H. T. F. Stoof, J. Phys, **B40**, 5629 (2008).
- 14. J. Geremia, J. Stockhom and H. Mubachi, Phys. Rev., A73, 042112 (2006).
- 15. K. D. Nelson, Xiao Li and D. S. Weiss, Nature Physics, 3, 566 (2007).
- 16. W. S. Bakr, J. I. Gilen, A. Peng, S. Folling and M. Greiner, Nature, 462, 7269 (2009).
- 17. Y. Lin, S. Jung, S. Maxwell, L. Turner, E. Tiesings and P. Lett, Phys. Rev. Lett. (PRL), **102**, 125301 (2009).
- 18. D. V. Vasilyev, I. V. Sokolov and E. S. Polzik, Phys. Rev., **A81**, 020302R (2010).
- 19. E. Zuethan, A. Gordeeka and A. Sorensen, Phys. Rev., **A84**, 043838 (2011).
- 20. K. Hammer, A. S. Sorensen and F. S. Poizik, Rev. Mod. Phys., **84**, 1041 (2011).

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